# Hartree-Fock methods and overview of material 

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## Overview of lectures

The aim of this series of lectures is to give you an overview of some selected post Hartree-Fock methods, with an empahsis on computational issues. The online material contains also material on second quantization and some background material on single-particle basis functions.

We will

- Start with a discussion of the Hartree-Fock (HF) approach since it serves as input to the other methods. This is our first lecture.
- The HF lecture will also allow us to bring in some mathematical aspects of basis functions. We will end the HF lecture with a discussion of a simple Python program which implements the HF algorithm.
- We move then over to Full Configuration Interaction (FCI) theory (popularly called the nuclear shell model). This is our second lecture and partly third lecture. FCI will also be covered in more detail by Maris' and Navratil's lectures.
- We discuss then approximative methods like many-body perturbation theory (MBPT), with links to FCI and applications to infinite nuclear matter. This is the third lecture.
- Finally we discuss coupled cluster (CC) theory with links to FCI and MBPT and applications to infinite nuclear matter. Coupled cluster theory are discussed in the fourth and fifth lectures. Numerical examples will be discussed and codes are fully available.

Other methods like Green's function approaches will be discussed by Soma and cluster methods by Neff.

## Why Hartree-Fock?

Hartree-Fock (HF) theory is an algorithm for finding an approximative expression for the ground state of a given Hamiltonian. The basic ingredients are

- Define a single-particle basis $\left\{\psi_{\alpha}\right\}$ so that

$$
\hat{h}^{\mathrm{HF}} \psi_{\alpha}=\varepsilon_{\alpha} \psi_{\alpha}
$$

with the Hartree-Fock Hamiltonian defined as

$$
\hat{h}^{\mathrm{HF}}=\hat{t}+\hat{u}_{\mathrm{ext}}+\hat{u}^{\mathrm{HF}}
$$

- The term $\hat{u}^{\mathrm{HF}}$ is a single-particle potential to be determined by the HF algorithm.
- The HF algorithm means to choose $\hat{u}^{\mathrm{HF}}$ in order to have

$$
\langle\hat{H}\rangle=E^{\mathrm{HF}}=\left\langle\Phi_{0}\right| \hat{H}\left|\Phi_{0}\right\rangle
$$

that is to find a local minimum with a Slater determinant $\Phi_{0}$ being the ansatz for the ground state.

- The variational principle ensures that $E^{\mathrm{HF}} \geq E_{0}$, with $E_{0}$ the exact ground state energy.


## Why Hartree-Fock?

We will show that the Hartree-Fock Hamiltonian $\hat{h}^{\mathrm{HF}}$ equals our definition of the operator $\hat{f}$ discussed in connection with the new definition of the normalordered Hamiltonian (see later lectures), that is we have, for a specific matrix element

$$
\langle p| \hat{h}^{\mathrm{HF}}|q\rangle=\langle p| \hat{f}|q\rangle=\langle p| \hat{t}+\hat{u}_{\mathrm{ext}}|q\rangle+\sum_{i \leq F}\langle p i| \hat{V}|q i\rangle_{A S}
$$

meaning that

$$
\langle p| \hat{u}^{\mathrm{HF}}|q\rangle=\sum_{i \leq F}\langle p i| \hat{V}|q i\rangle_{A S} .
$$

The so-called Hartree-Fock potential $\hat{u}^{\mathrm{HF}}$ brings an explicit medium dependence due to the summation over all single-particle states below the Fermi level $F$. It brings also in an explicit dependence on the two-body interaction (in nuclear physics we can also have complicated three- or higher-body forces). The twobody interaction, with its contribution from the other bystanding fermions, creates an effective mean field in which a given fermion moves, in addition to the external potential $\hat{u}_{\text {ext }}$ which confines the motion of the fermion. For systems like nuclei, there is no external confining potential. Nuclei are examples of self-bound systems, where the binding arises due to the intrinsic nature of the strong force. For nuclear systems thus, there would be no external one-body potential in the Hartree-Fock Hamiltonian.

## Definitions and notations

Before we proceed we need some definitions. We will assume that the interacting part of the Hamiltonian can be approximated by a two-body interaction. This means that our Hamiltonian is written as the sum of some onebody part and a twobody part

$$
\begin{equation*}
\hat{H}=\hat{H}_{0}+\hat{H}_{I}=\sum_{i=1}^{A} \hat{h}_{0}\left(x_{i}\right)+\sum_{i<j}^{A} \hat{v}\left(r_{i j}\right), \tag{1}
\end{equation*}
$$

with

$$
\begin{equation*}
H_{0}=\sum_{i=1}^{A} \hat{h}_{0}\left(x_{i}\right) . \tag{2}
\end{equation*}
$$

The onebody part $u_{\text {ext }}\left(x_{i}\right)$ is normally approximated by a harmonic oscillator or Woods-Saxon potential or for electronic systems the Coulomb interaction an electron feels from the nucleus. However, other potentials are fully possible, such as one derived from the self-consistent solution of the Hartree-Fock equations to be discussed here.

## Definitions and notations

Our Hamiltonian is invariant under the permutation (interchange) of two particles. Since we deal with fermions however, the total wave function is antisymmetric. Let $\hat{P}$ be an operator which interchanges two particles. Due to the symmetries we have ascribed to our Hamiltonian, this operator commutes with the total Hamiltonian,

$$
[\hat{H}, \hat{P}]=0
$$

meaning that $\Psi_{\lambda}\left(x_{1}, x_{2}, \ldots, x_{A}\right)$ is an eigenfunction of $\hat{P}$ as well, that is

$$
\hat{P}_{i j} \Psi_{\lambda}\left(x_{1}, x_{2}, \ldots, x_{i}, \ldots, x_{j}, \ldots, x_{A}\right)=\beta \Psi_{\lambda}\left(x_{1}, x_{2}, \ldots, x_{i}, \ldots, x_{j}, \ldots, x_{A}\right)
$$

where $\beta$ is the eigenvalue of $\hat{P}$. We have introduced the suffix $i j$ in order to indicate that we permute particles $i$ and $j$. The Pauli principle tells us that the total wave function for a system of fermions has to be antisymmetric, resulting in the eigenvalue $\beta=-1$.

## Definitions and notations

In our case we assume that we can approximate the exact eigenfunction with a Slater determinant

$$
\Phi\left(x_{1}, x_{2}, \ldots, x_{A}, \alpha, \beta, \ldots, \sigma\right)=\frac{1}{\sqrt{A!}}\left|\begin{array}{ccccc}
\psi_{\alpha}\left(x_{1}\right) & \psi_{\alpha}\left(x_{2}\right) & \ldots & \ldots & \psi_{\alpha}\left(x_{A}\right)  \tag{3}\\
\psi_{\beta}\left(x_{1}\right) & \psi_{\beta}\left(x_{2}\right) & \ldots & \ldots & \psi_{\beta}\left(x_{A}\right) \\
\ldots & \ldots & \ldots & \ldots & \ldots \\
\ldots & \ldots & \ldots & \ldots & \ldots \\
\psi_{\sigma}\left(x_{1}\right) & \psi_{\sigma}\left(x_{2}\right) & \ldots & \ldots & \psi_{\sigma}\left(x_{A}\right)
\end{array}\right|
$$

where $x_{i}$ stand for the coordinates and spin values of a particle $i$ and $\alpha, \beta, \ldots, \gamma$ are quantum numbers needed to describe remaining quantum numbers.

## Definitions and notations

The single-particle function $\psi_{\alpha}\left(x_{i}\right)$ are eigenfunctions of the onebody Hamiltonian $h_{i}$, that is

$$
\hat{h}_{0}\left(x_{i}\right)=\hat{t}\left(x_{i}\right)+\hat{u}_{\mathrm{ext}}\left(x_{i}\right),
$$

with eigenvalues

$$
\hat{h}_{0}\left(x_{i}\right) \psi_{\alpha}\left(x_{i}\right)=\left(\hat{t}\left(x_{i}\right)+\hat{u}_{\mathrm{ext}}\left(x_{i}\right)\right) \psi_{\alpha}\left(x_{i}\right)=\varepsilon_{\alpha} \psi_{\alpha}\left(x_{i}\right)
$$

The energies $\varepsilon_{\alpha}$ are the so-called non-interacting single-particle energies, or unperturbed energies. The total energy is in this case the sum over all singleparticle energies, if no two-body or more complicated many-body interactions are present.

## Definitions and notations

Let us denote the ground state energy by $E_{0}$. According to the variational principle we have

$$
E_{0} \leq E[\Phi]=\int \Phi^{*} \hat{H} \Phi d \tau
$$

where $\Phi$ is a trial function which we assume to be normalized

$$
\int \Phi^{*} \Phi d \tau=1
$$

where we have used the shorthand $d \tau=d x_{1} d r_{2} \ldots d r_{A}$.

## Brief reminder on some linear algebra properties

Before we proceed with a more compact representation of a Slater determinant, we would like to repeat some linear algebra properties which will be useful for our derivations of the energy as function of a Slater determinant, Hartree-Fock theory and later the nuclear shell model.

The inverse of a matrix is defined by

$$
\mathbf{A}^{-1} \cdot \mathbf{A}=I
$$

A unitary matrix $\mathbf{A}$ is one whose inverse is its adjoint

$$
\mathbf{A}^{-1}=\mathbf{A}^{\dagger}
$$

A real unitary matrix is called orthogonal and its inverse is equal to its transpose. A hermitian matrix is its own self-adjoint, that is

$$
\mathbf{A}=\mathbf{A}^{\dagger}
$$

## Basic Matrix Features

## Matrix Properties Reminder.

| Relations | Name | matrix elements |
| :---: | :---: | :---: |
| $A=A^{T}$ | symmetric | $a_{i j}=a_{j i}$ |
| $A=\left(A^{T}\right)^{-1}$ | real orthogonal | $\sum_{k} a_{i k} a_{j k}=\sum_{k} a_{k i} a_{k j}=\delta_{i j}$ |
| $A=A^{*}$ | real matrix | $a_{i j}=a_{i j}^{*}$ |
| $A=A^{\dagger}$ | hermitian | $a_{i j}=a_{j i}^{*}$ |
| $A=\left(A^{\dagger}\right)^{-1}$ | unitary | $\sum_{k} a_{i k} a_{j k}^{*}=\sum_{k} a_{k i}^{*} a_{k j}=\delta_{i j}$ |

## Basic Matrix Features

Since we will deal with Fermions (identical and indistinguishable particles) we will form an ansatz for a given state in terms of so-called Slater determinants determined by a chosen basis of single-particle functions.

For a given $n \times n$ matrix $\mathbf{A}$ we can write its determinant

$$
\operatorname{det}(\mathbf{A})=|\mathbf{A}|=\left|\begin{array}{ccccc}
a_{11} & a_{12} & \ldots & \ldots & a_{1 n} \\
a_{21} & a_{22} & \ldots & \ldots & a_{2 n} \\
\ldots & \ldots & \ldots & \ldots & \ldots \\
\ldots & \ldots & \ldots & \ldots & \ldots \\
a_{n 1} & a_{n 2} & \ldots & \ldots & a_{n n}
\end{array}\right|,
$$

in a more compact form as

$$
|\mathbf{A}|=\sum_{i=1}^{n!}(-1)^{p_{i}} \hat{P}_{i} a_{11} a_{22} \ldots a_{n n}
$$

where $\hat{P}_{i}$ is a permutation operator which permutes the column indices $1,2,3, \ldots, n$ and the sum runs over all $n$ ! permutations. The quantity $p_{i}$ represents the number of transpositions of column indices that are needed in order to bring a given permutation back to its initial ordering, in our case given by $a_{11} a_{22} \ldots a_{n n}$ here.

## Basic Matrix Features, simple $2 \times 2$ determinant

A simple $2 \times 2$ determinant illustrates this. We have

$$
\operatorname{det}(\mathbf{A})=\left|\begin{array}{ll}
a_{11} & a_{12} \\
a_{21} & a_{22}
\end{array}\right|=(-1)^{0} a_{11} a_{22}+(-1)^{1} a_{12} a_{21}
$$

where in the last term we have interchanged the column indices 1 and 2. The natural ordering we have chosen is $a_{11} a_{22}$.

## Definitions and notations

With the above we can rewrite our Slater determinant in a more compact form. In the Hartree-Fock method the trial function is the Slater determinant of Eq. (3) which can be rewritten as

$$
\Phi\left(x_{1}, x_{2}, \ldots, x_{A}, \alpha, \beta, \ldots, \nu\right)=\frac{1}{\sqrt{A!}} \sum_{P}(-)^{P} \hat{P} \psi_{\alpha}\left(x_{1}\right) \psi_{\beta}\left(x_{2}\right) \ldots \psi_{\nu}\left(x_{A}\right)=\sqrt{A!} \hat{A} \Phi_{H}
$$

where we have introduced the antisymmetrization operator $\hat{A}$ defined by the summation over all possible permutations of two particles.

## Definitions and notations

It is defined as

$$
\begin{equation*}
\hat{A}=\frac{1}{A!} \sum_{p}(-)^{p} \hat{P} \tag{4}
\end{equation*}
$$

with $p$ standing for the number of permutations. We have introduced for later use the so-called Hartree-function, defined by the simple product of all possible single-particle functions

$$
\Phi_{H}\left(x_{1}, x_{2}, \ldots, x_{A}, \alpha, \beta, \ldots, \nu\right)=\psi_{\alpha}\left(x_{1}\right) \psi_{\beta}\left(x_{2}\right) \ldots \psi_{\nu}\left(x_{A}\right)
$$

## Definitions and notations

Both $\hat{H}_{0}$ and $\hat{H}_{I}$ are invariant under all possible permutations of any two particles and hence commute with $\hat{A}$

$$
\begin{equation*}
\left[H_{0}, \hat{A}\right]=\left[H_{I}, \hat{A}\right]=0 \tag{5}
\end{equation*}
$$

Furthermore, $\hat{A}$ satisfies

$$
\begin{equation*}
\hat{A}^{2}=\hat{A}, \tag{6}
\end{equation*}
$$

since every permutation of the Slater determinant reproduces it.

## Definitions and notations

The expectation value of $\hat{H}_{0}$

$$
\int \Phi^{*} \hat{H}_{0} \Phi d \tau=A!\int \Phi_{H}^{*} \hat{A} \hat{H}_{0} \hat{A} \Phi_{H} d \tau
$$

is readily reduced to

$$
\int \Phi^{*} \hat{H}_{0} \Phi d \tau=A!\int \Phi_{H}^{*} \hat{H}_{0} \hat{A} \Phi_{H} d \tau
$$

where we have used Eqs. (5) and (6). The next step is to replace the antisymmetrization operator by its definition and to replace $\hat{H}_{0}$ with the sum of one-body operators

$$
\int \Phi^{*} \hat{H}_{0} \Phi d \tau=\sum_{i=1}^{A} \sum_{p}(-)^{p} \int \Phi_{H}^{*} \hat{h}_{0} \hat{P} \Phi_{H} d \tau
$$

## Definitions and notations

The integral vanishes if two or more particles are permuted in only one of the Hartree-functions $\Phi_{H}$ because the individual single-particle wave functions are orthogonal. We obtain then

$$
\int \Phi^{*} \hat{H}_{0} \Phi d \tau=\sum_{i=1}^{A} \int \Phi_{H}^{*} \hat{h}_{0} \Phi_{H} d \tau
$$

Orthogonality of the single-particle functions allows us to further simplify the integral, and we arrive at the following expression for the expectation values of the sum of one-body Hamiltonians

$$
\begin{equation*}
\int \Phi^{*} \hat{H}_{0} \Phi d \tau=\sum_{\mu=1}^{A} \int \psi_{\mu}^{*}(x) \hat{h}_{0} \psi_{\mu}(x) d x \tag{7}
\end{equation*}
$$

## Definitions and notations

We introduce the following shorthand for the above integral

$$
\langle\mu| \hat{h}_{0}|\mu\rangle=\int \psi_{\mu}^{*}(x) \hat{h}_{0} \psi_{\mu}(x) d x
$$

and rewrite Eq. (7) as

$$
\begin{equation*}
\int \Phi^{*} \hat{H}_{0} \Phi d \tau=\sum_{\mu=1}^{A}\langle\mu| \hat{h}_{0}|\mu\rangle \tag{8}
\end{equation*}
$$

## Definitions and notations

The expectation value of the two-body part of the Hamiltonian is obtained in a similar manner. We have

$$
\int \Phi^{*} \hat{H}_{I} \Phi d \tau=A!\int \Phi_{H}^{*} \hat{A} \hat{H}_{I} \hat{A} \Phi_{H} d \tau
$$

which reduces to

$$
\int \Phi^{*} \hat{H}_{I} \Phi d \tau=\sum_{i \leq j=1}^{A} \sum_{p}(-)^{p} \int \Phi_{H}^{*} \hat{v}\left(r_{i j}\right) \hat{P} \Phi_{H} d \tau
$$

by following the same arguments as for the one-body Hamiltonian.

## Definitions and notations

Because of the dependence on the inter-particle distance $r_{i j}$, permutations of any two particles no longer vanish, and we get

$$
\int \Phi^{*} \hat{H}_{I} \Phi d \tau=\sum_{i<j=1}^{A} \int \Phi_{H}^{*} \hat{v}\left(r_{i j}\right)\left(1-P_{i j}\right) \Phi_{H} d \tau
$$

where $P_{i j}$ is the permutation operator that interchanges particle $i$ and particle $j$. Again we use the assumption that the single-particle wave functions are orthogonal.

## Definitions and notations

We obtain

$$
\begin{align*}
\int \Phi^{*} \hat{H}_{I} \Phi d \tau=\frac{1}{2} \sum_{\mu=1}^{A} \sum_{\nu=1}^{A} & {\left[\int \psi_{\mu}^{*}\left(x_{i}\right) \psi_{\nu}^{*}\left(x_{j}\right) \hat{v}\left(r_{i j}\right) \psi_{\mu}\left(x_{i}\right) \psi_{\nu}\left(x_{j}\right) d x_{i} d x_{j}\right.}  \tag{9}\\
& \left.-\int \psi_{\mu}^{*}\left(x_{i}\right) \psi_{\nu}^{*}\left(x_{j}\right) \hat{v}\left(r_{i j}\right) \psi_{\nu}\left(x_{i}\right) \psi_{\mu}\left(x_{j}\right) d x_{i} d x_{j}\right] . \tag{10}
\end{align*}
$$

The first term is the so-called direct term. It gives rise to the Hartree term in Hartree-Fock theory, while the second is due to the Pauli principle and is called the exchange term and gives rise to the Fock term in the Hartree-Fock equations. The factor $1 / 2$ is introduced because we now run over all pairs twice.

## Definitions and notations

The last equation allows us to introduce some further definitions. The singleparticle wave functions $\psi_{\mu}(x)$, defined by the quantum numbers $\mu$ and $x$ are defined as the overlap

$$
\psi_{\mu}(x)=\langle x \mid \mu\rangle .
$$

## Definitions and notations

We introduce the following shorthands for the above two integrals

$$
\langle\mu \nu| \hat{v}|\mu \nu\rangle=\int \psi_{\mu}^{*}\left(x_{i}\right) \psi_{\nu}^{*}\left(x_{j}\right) \hat{v}\left(r_{i j}\right) \psi_{\mu}\left(x_{i}\right) \psi_{\nu}\left(x_{j}\right) d x_{i} d x_{j},
$$

and

$$
\langle\mu \nu| \hat{v}|\nu \mu\rangle=\int \psi_{\mu}^{*}\left(x_{i}\right) \psi_{\nu}^{*}\left(x_{j}\right) \hat{v}\left(r_{i j}\right) \psi_{\nu}\left(x_{i}\right) \psi_{\mu}\left(x_{j}\right) d x_{i} d x_{j}
$$

## Compact functional

Our functional can then be written in a compact version as

$$
E[\Phi]=\sum_{\mu}^{A}\langle\mu| \hat{h}_{0}|\mu\rangle+\frac{1}{2} \sum_{\mu \nu}^{A}[\langle\mu \nu| \hat{v}|\mu \nu\rangle-\langle\nu \mu| \hat{v}|\mu \nu\rangle] .
$$

## Properties of the interaction elements

Since the interaction is invariant under the interchange of two particles it means for example that we have

$$
\langle\mu \nu| \hat{v}|\mu \nu\rangle=\langle\nu \mu| \hat{v}|\nu \mu\rangle,
$$

or in the more general case

$$
\langle\mu \nu| \hat{v}|\sigma \tau\rangle=\langle\nu \mu| \hat{v}|\tau \sigma\rangle .
$$

## Redefining the matrix elements

The direct and exchange matrix elements can be brought together if we define the antisymmetrized matrix element

$$
\langle\mu \nu| \hat{v}|\mu \nu\rangle_{A S}=\langle\mu \nu| \hat{v}|\mu \nu\rangle-\langle\mu \nu| \hat{v}|\nu \mu\rangle,
$$

or for a general matrix element

$$
\langle\mu \nu| \hat{v}|\sigma \tau\rangle_{A S}=\langle\mu \nu| \hat{v}|\sigma \tau\rangle-\langle\mu \nu| \hat{v}|\tau \sigma\rangle .
$$

It has the symmetry property

$$
\langle\mu \nu| \hat{v}|\sigma \tau\rangle_{A S}=-\langle\mu \nu| \hat{v}|\tau \sigma\rangle_{A S}=-\langle\nu \mu| \hat{v}|\sigma \tau\rangle_{A S} .
$$

The antisymmetric matrix element is also hermitian, implying

$$
\langle\mu \nu| \hat{v}|\sigma \tau\rangle_{A S}=\langle\sigma \tau| \hat{v}|\mu \nu\rangle_{A S} .
$$

## Rewriting the energy functional

With these notations we rewrite the energy functional as

$$
\begin{equation*}
\int \Phi^{*} \hat{H}_{I} \Phi d \tau=\frac{1}{2} \sum_{\mu=1}^{A} \sum_{\nu=1}^{A}\langle\mu \nu| \hat{v}|\mu \nu\rangle_{A S} . \tag{11}
\end{equation*}
$$

Adding the contribution from the one-body operator $\hat{H}_{0}$ to (11) we obtain the energy functional

$$
\begin{equation*}
E[\Phi]=\sum_{\mu=1}^{A}\langle\mu| h|\mu\rangle+\frac{1}{2} \sum_{\mu=1}^{A} \sum_{\nu=1}^{A}\langle\mu \nu| \hat{v}|\mu \nu\rangle_{A S} . \tag{12}
\end{equation*}
$$

In our coordinate space derivations below we will spell out the Hartree-Fock equations in terms of their integrals.

## Reminder on Variational Calculus and Lagrangian Multipliers

The calculus of variations involves problems where the quantity to be minimized or maximized is an integral.

In the general case we have an integral of the type

$$
E[\Phi]=\int_{a}^{b} f\left(\Phi(x), \frac{\partial \Phi}{\partial x}, x\right) d x
$$

where $E$ is the quantity which is sought minimized or maximized. The problem is that although $f$ is a function of the variables $\Phi, \partial \Phi / \partial x$ and $x$, the exact dependence of $\Phi$ on $x$ is not known. This means again that even though the integral has fixed limits $a$ and $b$, the path of integration is not known. In our case the unknown quantities are the single-particle wave functions and we wish to choose an integration path which makes the functional $E[\Phi]$ stationary. This means that we want to find minima, or maxima or saddle points. In physics we search normally for minima. Our task is therefore to find the minimum of $E[\Phi]$ so that its variation $\delta E$ is zero subject to specific constraints. In our case the constraints appear as the integral which expresses the orthogonality of the single-particle wave functions. The constraints can be treated via the technique of Lagrangian multipliers

## Variational Calculus and Lagrangian Multipliers, simple example

Let us specialize to the expectation value of the energy for one particle in three-dimensions. This expectation value reads

$$
E=\int d x d y d z \psi^{*}(x, y, z) \hat{H} \psi(x, y, z)
$$

with the constraint

$$
\int d x d y d z \psi^{*}(x, y, z) \psi(x, y, z)=1
$$

and a Hamiltonian

$$
\hat{H}=-\frac{1}{2} \nabla^{2}+V(x, y, z)
$$

We will, for the sake of notational convenience, skip the variables $x, y, z$ below, and write for example $V(x, y, z)=V$.

## Manipulating terms

The integral involving the kinetic energy can be written as, with the function $\psi$ vanishing strongly for large values of $x, y, z$ (given here by the limits $a$ and $b$ ),

$$
\int_{a}^{b} d x d y d z \psi^{*}\left(-\frac{1}{2} \nabla^{2}\right) \psi d x d y d z=\left.\psi^{*} \nabla \psi\right|_{a} ^{b}+\int_{a}^{b} d x d y d z \frac{1}{2} \nabla \psi^{*} \nabla \psi
$$

We will drop the limits $a$ and $b$ in the remaining discussion. Inserting this expression into the expectation value for the energy and taking the variational minimum we obtain

$$
\delta E=\delta\left\{\int d x d y d z\left(\frac{1}{2} \nabla \psi^{*} \nabla \psi+V \psi^{*} \psi\right)\right\}=0 .
$$

## Adding the Lagrangian multiplier

The constraint appears in integral form as

$$
\int d x d y d z \psi^{*} \psi=\text { constant }
$$

and multiplying with a Lagrangian multiplier $\lambda$ and taking the variational minimum we obtain the final variational equation

$$
\delta\left\{\int d x d y d z\left(\frac{1}{2} \nabla \psi^{*} \nabla \psi+V \psi^{*} \psi-\lambda \psi^{*} \psi\right)\right\}=0
$$

We introduce the function $f$

$$
f=\frac{1}{2} \nabla \psi^{*} \nabla \psi+V \psi^{*} \psi-\lambda \psi^{*} \psi=\frac{1}{2}\left(\psi_{x}^{*} \psi_{x}+\psi_{y}^{*} \psi_{y}+\psi_{z}^{*} \psi_{z}\right)+V \psi^{*} \psi-\lambda \psi^{*} \psi
$$

where we have skipped the dependence on $x, y, z$ and introduced the shorthand $\psi_{x}, \psi_{y}$ and $\psi_{z}$ for the various derivatives.

## And with the Euler-Lagrange equations we get

For $\psi^{*}$ the Euler-Lagrange equations yield

$$
\frac{\partial f}{\partial \psi^{*}}-\frac{\partial}{\partial x} \frac{\partial f}{\partial \psi_{x}^{*}}-\frac{\partial}{\partial y} \frac{\partial f}{\partial \psi_{y}^{*}}-\frac{\partial}{\partial z} \frac{\partial f}{\partial \psi_{z}^{*}}=0
$$

which results in

$$
-\frac{1}{2}\left(\psi_{x x}+\psi_{y y}+\psi_{z z}\right)+V \psi=\lambda \psi .
$$

We can then identify the Lagrangian multiplier as the energy of the system. The last equation is nothing but the standard Schroedinger equation and the variational approach discussed here provides a powerful method for obtaining approximate solutions of the wave function.

## Hartree-Fock by varying the coefficients of a wave function expansion

In deriving the Hartree-Fock equations, we will expand the single-particle functions in a known basis and vary the coefficients, that is, the new singleparticle wave function is written as a linear expansion in terms of a fixed chosen
orthogonal basis (for example the well-known harmonic oscillator functions or the hydrogen-like functions etc). We define our new Hartree-Fock single-particle basis by performing a unitary transformation on our previous basis (labelled with greek indices) as

$$
\begin{equation*}
\psi_{p}^{H F}=\sum_{\lambda} C_{p \lambda} \phi_{\lambda} \tag{13}
\end{equation*}
$$

In this case we vary the coefficients $C_{p \lambda}$. If the basis has infinitely many solutions, we need to truncate the above sum. We assume that the basis $\phi_{\lambda}$ is orthogonal. A unitary transformation keeps the orthogonality, as discussed in exercise 1 below.

## More on linear algebra

In the previous slide we stated that a unitary transformation keeps the orthogonality, as discussed in exercise 1 below. To see this consider first a basis of vectors $\mathbf{v}_{i}$,

$$
\mathbf{v}_{i}=\left[\begin{array}{c}
v_{i 1} \\
\cdots \\
\cdots \\
v_{i n}
\end{array}\right]
$$

We assume that the basis is orthogonal, that is

$$
\mathbf{v}_{j}^{T} \mathbf{v}_{i}=\delta_{i j}
$$

An orthogonal or unitary transformation

$$
\mathbf{w}_{i}=\mathbf{U v}_{i}
$$

preserves the dot product and orthogonality since

$$
\mathbf{w}_{j}^{T} \mathbf{w}_{i}=\left(\mathbf{U} \mathbf{v}_{j}\right)^{T} \mathbf{U} \mathbf{v}_{i}=\mathbf{v}_{j}^{T} \mathbf{U}^{T} \mathbf{U} \mathbf{v}_{i}=\mathbf{v}_{j}^{T} \mathbf{v}_{i}=\delta_{i j}
$$

## Coefficients of a wave function expansion

This means that if the coefficients $C_{p \lambda}$ belong to a unitary or orthogonal trasformation (using the Dirac bra-ket notation)

$$
|p\rangle=\sum_{\lambda} C_{p \lambda}|\lambda\rangle,
$$

orthogonality is preserved, that is $\langle\alpha \mid \beta\rangle=\delta_{\alpha \beta}$ and $\langle p \mid q\rangle=\delta_{p q}$.
This propertry is extremely useful when we build up a basis of many-body Stater determinant based states.

Note also that although a basis $|\alpha\rangle$ contains an infinity of states, for practical calculations we have always to make some truncations.

## More Basic Matrix Features, simple $2 \times 2$ determinant, useful property of determinants

Before we develop the Hartree-Fock equations, there is another very useful property of determinants that we will use both in connection with Hartree-Fock calculations and later shell-model calculations.

Consider the following determinant

$$
\left|\begin{array}{cc}
\alpha_{1} b_{11}+\alpha_{2} s b_{12} & a_{12} \\
\alpha_{1} b_{21}+\alpha_{2} b_{22} & a_{22}
\end{array}\right|=\alpha_{1}\left|\begin{array}{ll}
b_{11} & a_{12} \\
b_{21} & a_{22}
\end{array}\right|+\alpha_{2}\left|\begin{array}{ll}
b_{12} & a_{12} \\
b_{22} & a_{22}
\end{array}\right|
$$

## More Basic Matrix Features, $n \times n$ determinant

We can generalize this to an $n \times n$ matrix and have

$$
\left|\begin{array}{cccccc}
a_{11} & a_{12} & \ldots & \sum_{k=1}^{n} c_{k} b_{1 k} & \ldots & a_{1 n} \\
a_{21} & a_{22} & \ldots & \sum_{k=1}^{n} c_{k} b_{2 k} & \ldots & a_{2 n} \\
\ldots & \ldots & \ldots & \ldots & \ldots & \ldots \\
\ldots & \ldots & \ldots & \ldots & \ldots & \ldots \\
a_{n 1} & a_{n 2} & \ldots & \sum_{k=1}^{n} c_{k} b_{n k} & \ldots & a_{n n}
\end{array}\right|=\sum_{k=1}^{n} c_{k}\left|\begin{array}{cccccc}
a_{11} & a_{12} & \ldots & b_{1 k} & \ldots & a_{1 n} \\
a_{21} & a_{22} & \ldots & b_{2 k} & \ldots & a_{2 n} \\
\ldots & \ldots & \ldots & \ldots & \ldots & \ldots \\
\ldots & \ldots & \ldots & \ldots & \ldots & \ldots \\
a_{n 1} & a_{n 2} & \ldots & b_{n k} & \ldots & a_{n n}
\end{array}\right|
$$

This is a property we will use in our Hartree-Fock discussions.

## More Basic Matrix Features, a general $n \times n$ determinant

We can generalize the previous results, now with all elements $a_{i j}$ being given as functions of linear combinations of various coefficients $c$ and elements $b_{i j}$,

$$
\left|\begin{array}{cccccc}
\sum_{k=1}^{n} b_{1 k} c_{k 1} & \sum_{k=1}^{n} b_{1 k} c_{k 2} & \ldots & \sum_{k=1}^{n} b_{1 k} c_{k j} & \ldots & \sum_{k=1}^{n} b_{1 k} c_{k n} \\
\sum_{k=1}^{n} b_{2 k} c_{k 1} & \sum_{k=1}^{n} b_{2 k} c_{k 2} & \ldots & \sum_{k=1}^{n=1} b_{2 k} c_{k j} & \ldots & \sum_{k=1}^{n} b_{2 k} c_{k n} \\
\ldots & \cdots & \ldots & \cdots & \ldots & \cdots \\
\cdots & \cdots & \cdots & \cdots & \cdots & \cdots \\
\sum_{k=1}^{n} b_{n k} c_{k 1} & \sum_{k=1}^{n} b_{n k} c_{k 2} & \cdots & \sum_{k=1}^{n} b_{n k} c_{k j} & \cdots & \sum_{k=1}^{n} b_{n k} c_{k n}
\end{array}\right|=\operatorname{det}(\mathbf{C}) \operatorname{det}(\mathbf{B}),
$$

where $\operatorname{det}(\mathbf{C})$ and $\operatorname{det}(\mathbf{B})$ are the determinants of $n \times n$ matrices with elements $c_{i j}$ and $b_{i j}$ respectively. This is a property we will use in our Hartree-Fock discussions. Convince yourself about the correctness of the above expression by setting $n=2$.

## A general Slater determinant

With our definition of the new basis in terms of an orthogonal basis we have

$$
\psi_{p}(x)=\sum_{\lambda} C_{p \lambda} \phi_{\lambda}(x)
$$

If the coefficients $C_{p \lambda}$ belong to an orthogonal or unitary matrix, the new basis is also orthogonal. Our Slater determinant in the new basis $\psi_{p}(x)$ is written as
which is nothing but $\operatorname{det}(\mathbf{C}) \operatorname{det}(\Phi)$, with $\operatorname{det}(\Phi)$ being the determinant given by the basis functions $\phi_{\lambda}(x)$.

## Hartree-Fock by varying the coefficients of a wave function expansion

It is normal to choose a single-particle basis defined as the eigenfunctions of parts of the full Hamiltonian. The typical situation consists of the solutions of the one-body part of the Hamiltonian, that is we have

$$
\hat{h}_{0} \phi_{\lambda}=\epsilon_{\lambda} \phi_{\lambda}
$$

The single-particle wave functions $\phi_{\lambda}(\mathbf{r})$, defined by the quantum numbers $\lambda$ and $\mathbf{r}$ are defined as the overlap

$$
\phi_{\lambda}(\mathbf{r})=\langle\mathbf{r} \mid \lambda\rangle .
$$

## Hartree-Fock by varying the coefficients of a wave function expansion

In our discussions hereafter we will use our definitions of single-particle states above and below the Fermi $(F)$ level given by the labels $i j k l \cdots \leq F$ for so-called single-hole states and $a b c d \cdots>F$ for so-called particle states. For general single-particle states we employ the labels pqrs....

## Hartree-Fock by varying the coefficients of a wave function expansion

In Eq. (12), restated here

$$
E[\Phi]=\sum_{\mu=1}^{A}\langle\mu| h|\mu\rangle+\frac{1}{2} \sum_{\mu=1}^{A} \sum_{\nu=1}^{A}\langle\mu \nu| \hat{v}|\mu \nu\rangle_{A S},
$$

we found the expression for the energy functional in terms of the basis function $\phi_{\lambda}(\mathbf{r})$. We then varied the above energy functional with respect to the basis functions $|\mu\rangle$. Now we are interested in defining a new basis defined in terms of
a chosen basis as defined in Eq. (13). We can then rewrite the energy functional as

$$
\begin{equation*}
E\left[\Phi^{H F}\right]=\sum_{i=1}^{A}\langle i| h|i\rangle+\frac{1}{2} \sum_{i j=1}^{A}\langle i j| \hat{v}|i j\rangle_{A S}, \tag{14}
\end{equation*}
$$

where $\Phi^{H F}$ is the new Slater determinant defined by the new basis of Eq. (13).

## Hartree-Fock by varying the coefficients of a wave function expansion

Using Eq. (13) we can rewrite Eq. (14) as

$$
\begin{equation*}
E[\Psi]=\sum_{i=1}^{A} \sum_{\alpha \beta} C_{i \alpha}^{*} C_{i \beta}\langle\alpha| h|\beta\rangle+\frac{1}{2} \sum_{i j=1}^{A} \sum_{\alpha \beta \gamma \delta} C_{i \alpha}^{*} C_{j \beta}^{*} C_{i \gamma} C_{j \delta}\langle\alpha \beta| \hat{v}|\gamma \delta\rangle_{A S} . \tag{15}
\end{equation*}
$$

## Hartree-Fock by varying the coefficients of a wave function expansion

We wish now to minimize the above functional. We introduce again a set of Lagrange multipliers, noting that since $\langle i \mid j\rangle=\delta_{i, j}$ and $\langle\alpha \mid \beta\rangle=\delta_{\alpha, \beta}$, the coefficients $C_{i \gamma}$ obey the relation

$$
\langle i \mid j\rangle=\delta_{i, j}=\sum_{\alpha \beta} C_{i \alpha}^{*} C_{i \beta}\langle\alpha \mid \beta\rangle=\sum_{\alpha} C_{i \alpha}^{*} C_{i \alpha}
$$

which allows us to define a functional to be minimized that reads

$$
\begin{equation*}
F\left[\Phi^{H F}\right]=E\left[\Phi^{H F}\right]-\sum_{i=1}^{A} \epsilon_{i} \sum_{\alpha} C_{i \alpha}^{*} C_{i \alpha} \tag{16}
\end{equation*}
$$

## Hartree-Fock by varying the coefficients of a wave function expansion

Minimizing with respect to $C_{i \alpha}^{*}$, remembering that the equations for $C_{i \alpha}^{*}$ and $C_{i \alpha}$ can be written as two independent equations, we obtain

$$
\frac{d}{d C_{i \alpha}^{*}}\left[E\left[\Phi^{H F}\right]-\sum_{j} \epsilon_{j} \sum_{\alpha} C_{j \alpha}^{*} C_{j \alpha}\right]=0
$$

which yields for every single-particle state $i$ and index $\alpha$ (recalling that the coefficients $C_{i \alpha}$ are matrix elements of a unitary (or orthogonal for a real symmetric matrix) matrix) the following Hartree-Fock equations

$$
\sum_{\beta} C_{i \beta}\langle\alpha| h|\beta\rangle+\sum_{j=1}^{A} \sum_{\beta \gamma \delta} C_{j \beta}^{*} C_{j \delta} C_{i \gamma}\langle\alpha \beta| \hat{v}|\gamma \delta\rangle_{A S}=\epsilon_{i}^{H F} C_{i \alpha}
$$

## Hartree-Fock by varying the coefficients of a wave function expansion

We can rewrite this equation as (changing dummy variables)

$$
\sum_{\beta}\left\{\langle\alpha| h|\beta\rangle+\sum_{j}^{A} \sum_{\gamma \delta} C_{j \gamma}^{*} C_{j \delta}\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S}\right\} C_{i \beta}=\epsilon_{i}^{H F} C_{i \alpha} .
$$

Note that the sums over greek indices run over the number of basis set functions (in principle an infinite number).

## Hartree-Fock by varying the coefficients of a wave function expansion

Defining

$$
h_{\alpha \beta}^{H F}=\langle\alpha| h|\beta\rangle+\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{*} C_{j \delta}\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S}
$$

we can rewrite the new equations as

$$
\begin{equation*}
\sum_{\gamma} h_{\alpha \beta}^{H F} C_{i \beta}=\epsilon_{i}^{H F} C_{i \alpha} . \tag{17}
\end{equation*}
$$

The latter is nothing but a standard eigenvalue problem.
It suffices to tabulate the matrix elements $\langle\alpha| h|\beta\rangle$ and $\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S}$ once and for all. Successive iterations require thus only a look-up in tables over one-body and two-body matrix elements. These details will be discussed below when we solve the Hartree-Fock equations numerically.

## Hartree-Fock algorithm

Our Hartree-Fock matrix is thus

$$
\hat{h}_{\alpha \beta}^{H F}=\langle\alpha| \hat{h}_{0}|\beta\rangle+\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{*} C_{j \delta}\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S}
$$

The Hartree-Fock equations are solved in an iterative waym starting with a guess for the coefficients $C_{j \gamma}=\delta_{j, \gamma}$ and solving the equations by diagonalization till the new single-particle energies $\epsilon_{i}^{\mathrm{HF}}$ do not change anymore by a prefixed quantity.

## Hartree-Fock algorithm

Normally we assume that the single-particle basis $|\beta\rangle$ forms an eigenbasis for the operator $\hat{h}_{0}$, meaning that the Hartree-Fock matrix becomes

$$
\hat{h}_{\alpha \beta}^{H F}=\epsilon_{\alpha} \delta_{\alpha, \beta}+\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{*} C_{j \delta}\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S}
$$

The Hartree-Fock eigenvalue problem

$$
\sum_{\beta} \hat{h}_{\alpha \beta}^{H F} C_{i \beta}=\epsilon_{i}^{\mathrm{HF}} C_{i \alpha},
$$

can be written out in a more compact form as

$$
\hat{h}^{H F} \hat{C}=\epsilon^{\mathrm{HF}} \hat{C}
$$

## Hartree-Fock algorithm

The Hartree-Fock equations are, in their simplest form, solved in an iterative way, starting with a guess for the coefficients $C_{i \alpha}$. We label the coefficients as $C_{i \alpha}^{(n)}$, where the subscript $n$ stands for iteration $n$. To set up the algorithm we can proceed as follows:

- We start with a guess $C_{i \alpha}^{(0)}=\delta_{i, \alpha}$. Alternatively, we could have used random starting values as long as the vectors are normalized. Another possibility is to give states below the Fermi level a larger weight.
- The Hartree-Fock matrix simplifies then to (assuming that the coefficients $C_{i \alpha}$ are real)

$$
\hat{h}_{\alpha \beta}^{H F}=\epsilon_{\alpha} \delta_{\alpha, \beta}+\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{(0)} C_{j \delta}^{(0)}\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S} .
$$

## Hartree-Fock algorithm

Solving the Hartree-Fock eigenvalue problem yields then new eigenvectors $C_{i \alpha}^{(1)}$ and eigenvalues $\epsilon_{i}^{H F(1)}$.

- With the new eigenvalues we can set up a new Hartree-Fock potential

$$
\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{(1)} C_{j \delta}^{(1)}\langle\alpha \gamma| \hat{v}|\beta \delta\rangle_{A S}
$$

The diagonalization with the new Hartree-Fock potential yields new eigenvectors and eigenvalues. This process is continued till for example

$$
\frac{\sum_{p}\left|\epsilon_{i}^{(n)}-\epsilon_{i}^{(n-1)}\right|}{m} \leq \lambda,
$$

where $\lambda$ is a user prefixed quantity ( $\lambda \sim 10^{-8}$ or smaller) and $p$ runs over all calculated single-particle energies and $m$ is the number of single-particle states.

## Analysis of Hartree-Fock equations and Koopman's theorem

We can rewrite the ground state energy by adding and subtracting $\hat{u}^{H F}\left(x_{i}\right)$

$$
E_{0}^{H F}=\left\langle\Phi_{0}\right| \hat{H}\left|\Phi_{0}\right\rangle=\sum_{i \leq F}^{A}\langle i| \hat{h}_{0}+\hat{u}^{H F}|j\rangle+\frac{1}{2} \sum_{i \leq F}^{A} \sum_{j \leq F}^{A}[\langle i j| \hat{v}|i j\rangle-\langle i j| \hat{v}|j i\rangle]-\sum_{i \leq F}^{A}\langle i| \hat{u}^{H F}|i\rangle,
$$

which results in

$$
E_{0}^{H F}=\sum_{i \leq F}^{A} \varepsilon_{i}^{H F}+\frac{1}{2} \sum_{i \leq F}^{A} \sum_{j \leq F}^{A}[\langle i j| \hat{v}|i j\rangle-\langle i j| \hat{v}|j i\rangle]-\sum_{i \leq F}^{A}\langle i| \hat{u}^{H F}|i\rangle .
$$

Our single-particle states $i j k \ldots$ are now single-particle states obtained from the solution of the Hartree-Fock equations.

## Analysis of Hartree-Fock equations and Koopman's theorem

Using our definition of the Hartree-Fock single-particle energies we obtain then the following expression for the total ground-state energy

$$
E_{0}^{H F}=\sum_{i \leq F}^{A} \varepsilon_{i}-\frac{1}{2} \sum_{i \leq F}^{A} \sum_{j \leq F}^{A}[\langle i j| \hat{v}|i j\rangle-\langle i j| \hat{v}|j i\rangle] .
$$

This form will be used in our discussion of Koopman's theorem.

## Analysis of Hartree-Fock equations and Koopman's theorem

Atomic physics case. We have

$$
E\left[\Phi^{\mathrm{HF}}(N)\right]=\sum_{i=1}^{H}\langle i| \hat{h}_{0}|i\rangle+\frac{1}{2} \sum_{i j=1}^{N}\langle i j| \hat{v}|i j\rangle_{A S},
$$

where $\Phi^{\mathrm{HF}}(N)$ is the new Slater determinant defined by the new basis of Eq. (13) for $N$ electrons (same $Z$ ). If we assume that the single-particle wave functions in the new basis do not change when we remove one electron or add one electron, we can then define the corresponding energy for the $N-1$ systems as

$$
E\left[\Phi^{\mathrm{HF}}(N-1)\right]=\sum_{i=1 ; i \neq k}^{N}\langle i| \hat{h}_{0}|i\rangle+\frac{1}{2} \sum_{i j=1 ; i, j \neq k}^{N}\langle i j| \hat{v}|i j\rangle_{A S},
$$

where we have removed a single-particle state $k \leq F$, that is a state below the Fermi level.

## Analysis of Hartree-Fock equations and Koopman's theorem

Calculating the difference

$$
E\left[\Phi^{\mathrm{HF}}(N)\right]-E\left[\Phi^{\mathrm{HF}}(N-1)\right]=\langle k| \hat{h}_{0}|k\rangle+\frac{1}{2} \sum_{i=1 ; i \neq k}^{N}\langle i k| \hat{v}|i k\rangle_{A S} \frac{1}{2} \sum_{j=1 ; j \neq k}^{N}\langle k j| \hat{v}|k j\rangle_{A S},
$$

we obtain

$$
E\left[\Phi^{\mathrm{HF}}(N)\right]-E\left[\Phi^{\mathrm{HF}}(N-1)\right]=\langle k| \hat{h}_{0}|k\rangle+\frac{1}{2} \sum_{j=1}^{N}\langle k j| \hat{v}|k j\rangle_{A S}
$$

which is just our definition of the Hartree-Fock single-particle energy

$$
E\left[\Phi^{\mathrm{HF}}(N)\right]-E\left[\Phi^{\mathrm{HF}}(N-1)\right]=\epsilon_{k}^{\mathrm{HF}}
$$

## Analysis of Hartree-Fock equations and Koopman's theorem

Similarly, we can now compute the difference (we label the single-particle states above the Fermi level as $a b c d>F$ )

$$
E\left[\Phi^{\mathrm{HF}}(N+1)\right]-E\left[\Phi^{\mathrm{HF}}(N)\right]=\epsilon_{a}^{\mathrm{HF}} .
$$

These two equations can thus be used to the electron affinity or ionization energies, respectively. Koopman's theorem states that for example the ionization energy of a closed-shell system is given by the energy of the highest occupied single-particle state. If we assume that changing the number of electrons from $N$ to $N+1$ does not change the Hartree-Fock single-particle energies and eigenfunctions, then Koopman's theorem simply states that the ionization energy of an atom is given by the single-particle energy of the last bound state. In a similar way, we can also define the electron affinities.

## Analysis of Hartree-Fock equations and Koopman's theorem

As an example, consider a simple model for atomic sodium, Na. Neutral sodium has eleven electrons, with the weakest bound one being confined the 3 s single-particle quantum numbers. The energy needed to remove an electron from neutral sodium is rather small, 5.1391 eV , a feature which pertains to all alkali metals. Having performed a Hartree-Fock calculation for neutral sodium would then allows us to compute the ionization energy by using the single-particle energy for the $3 s$ states, namely $\epsilon_{3 s}^{\mathrm{HF}}$.

From these considerations, we see that Hartree-Fock theory allows us to make a connection between experimental observables (here ionization and affinity
energies) and the underlying interactions between particles. In this sense, we are now linking the dynamics and structure of a many-body system with the laws of motion which govern the system. Our approach is a reductionistic one, meaning that we want to understand the laws of motion in terms of the particles or degrees of freedom which we believe are the fundamental ones. Our Slater determinant, being constructed as the product of various single-particle functions, follows this philosophy.

## Analysis of Hartree-Fock equations, Koopman's theorem

With similar arguments as in atomic physics, we can now use Hartree-Fock theory to make a link between nuclear forces and separation energies. Changing to nuclear system, we define

$$
E\left[\Phi^{\mathrm{HF}}(A)\right]=\sum_{i=1}^{A}\langle i| \hat{h}_{0}|i\rangle+\frac{1}{2} \sum_{i j=1}^{A}\langle i j| \hat{v}|i j\rangle_{A S},
$$

where $\Phi^{\mathrm{HF}}(A)$ is the new Slater determinant defined by the new basis of Eq. (13) for $A$ nucleons, where $A=N+Z$, with $N$ now being the number of neutrons and $Z$ th enumber of protons. If we assume again that the single-particle wave functions in the new basis do not change from a nucleus with $A$ nucleons to a nucleus with $A-1$ nucleons, we can then define the corresponding energy for the $A-1$ systems as

$$
E\left[\Phi^{\mathrm{HF}}(A-1)\right]=\sum_{i=1 ; i \neq k}^{A}\langle i| \hat{h}_{0}|i\rangle+\frac{1}{2} \sum_{i j=1 ; i, j \neq k}^{A}\langle i j| \hat{v}|i j\rangle_{A S},
$$

where we have removed a single-particle state $k \leq F$, that is a state below the Fermi level.

## Analysis of Hartree-Fock equations and Koopman's theorem

Calculating the difference

$$
E\left[\Phi^{\mathrm{HF}}(A)\right]-E\left[\Phi^{\mathrm{HF}}(A-1)\right]=\langle k| \hat{h}_{0}|k\rangle+\frac{1}{2} \sum_{i=1 ; i \neq k}^{A}\langle i k| \hat{v}|i k\rangle_{A S} \frac{1}{2} \sum_{j=1 ; j \neq k}^{A}\langle k j| \hat{v}|k j\rangle_{A S}
$$

which becomes

$$
E\left[\Phi^{\mathrm{HF}}(A)\right]-E\left[\Phi^{\mathrm{HF}}(A-1)\right]=\langle k| \hat{h}_{0}|k\rangle+\frac{1}{2} \sum_{j=1}^{A}\langle k j| \hat{v}|k j\rangle_{A S}
$$

which is just our definition of the Hartree-Fock single-particle energy

$$
E\left[\Phi^{\mathrm{HF}}(A)\right]-E\left[\Phi^{\mathrm{HF}}(A-1)\right]=\epsilon_{k}^{\mathrm{HF}}
$$

## Analysis of Hartree-Fock equations and Koopman's theorem

Similarly, we can now compute the difference (recall that the single-particle states $a b c d>F$ )

$$
E\left[\Phi^{\mathrm{HF}}(A+1)\right]-E\left[\Phi^{\mathrm{HF}}(A)\right]=\epsilon_{a}^{\mathrm{HF}}
$$

If we then recall that the binding energy differences

$$
B E(A)-B E(A-1) \quad \text { and } \quad B E(A+1)-B E(A),
$$

define the separation energies, we see that the Hartree-Fock single-particle energies can be used to define separation energies. We have thus our first link between nuclear forces (included in the potential energy term) and an observable quantity defined by differences in binding energies.

## Analysis of Hartree-Fock equations and Koopman's theorem

We have thus the following interpretations (if the single-particle field do not change)

$$
B E(A)-B E(A-1) \approx E\left[\Phi^{\mathrm{HF}}(A)\right]-E\left[\Phi^{\mathrm{HF}}(A-1)\right]=\epsilon_{k}^{\mathrm{HF}},
$$

and

$$
B E(A+1)-B E(A) \approx E\left[\Phi^{\mathrm{HF}}(A+1)\right]-E\left[\Phi^{\mathrm{HF}}(A)\right]=\epsilon_{a}^{\mathrm{HF}} .
$$

If we use ${ }^{16} \mathrm{O}$ as our closed-shell nucleus, we could then interpret the separation energy

$$
B E\left({ }^{16} \mathrm{O}\right)-B E\left({ }^{15} \mathrm{O}\right) \approx \epsilon_{0 p_{1 / 2}^{\nu}}^{\mathrm{HF}}
$$

and

$$
B E\left({ }^{16} \mathrm{O}\right)-B E\left({ }^{15} \mathrm{~N}\right) \approx \epsilon_{0 p_{1 / 2}^{\pi}}^{\mathrm{HF}}
$$

## Analysis of Hartree-Fock equations and Koopman's theorem

Similalry, we could interpret

$$
B E\left({ }^{17} \mathrm{O}\right)-B E\left({ }^{16} \mathrm{O}\right) \approx \epsilon_{0 d_{5 / 2}^{\nu}}^{\mathrm{HF}}
$$

and

$$
B E\left({ }^{17} \mathrm{~F}\right)-B E\left({ }^{16} \mathrm{O}\right) \approx \epsilon_{0 d_{5 / 2}^{\pi}}^{\mathrm{HF}}
$$

We can continue like this for all $A \pm 1$ nuclei where $A$ is a good closed-shell (or subshell closure) nucleus. Examples are ${ }^{22} \mathrm{O},{ }^{24} \mathrm{O},{ }^{40} \mathrm{Ca},{ }^{48} \mathrm{Ca},{ }^{52} \mathrm{Ca},{ }^{54} \mathrm{Ca},{ }^{56} \mathrm{Ni}$, ${ }^{68} \mathrm{Ni},{ }^{78} \mathrm{Ni},{ }^{90} \mathrm{Zr},{ }^{88} \mathrm{Sr},{ }^{100} \mathrm{Sn},{ }^{132} \mathrm{Sn}$ and ${ }^{208} \mathrm{~Pb}$, to mention some possile cases.

## Analysis of Hartree-Fock equations and Koopman's theorem

We can thus make our first interpretation of the separation energies in terms of the simplest possible many-body theory. If we also recall that the so-called energy gap for neutrons (or protons) is defined as

$$
\Delta S_{n}=2 B E(N, Z)-B E(N-1, Z)-B E(N+1, Z),
$$

for neutrons and the corresponding gap for protons

$$
\Delta S_{p}=2 B E(N, Z)-B E(N, Z-1)-B E(N, Z+1),
$$

we can define the neutron and proton energy gaps for ${ }^{16} \mathrm{O}$ as

$$
\Delta S_{\nu}=\epsilon_{0 d_{5 / 2}}^{\mathrm{HF}}-\epsilon_{0 p_{1 / 2}}^{\mathrm{HF}},
$$

and

$$
\Delta S_{\pi}=\epsilon_{0 d_{5 / 2}^{\pi}}^{\mathrm{HF}}-\epsilon_{\epsilon_{p_{1 / 2}^{\pi}}^{\mathrm{HF}}} .
$$

* 

Exercise 1: Hartree-Fock Slater determinant
Consider a Slater determinant built up of orthogonal single-particle orbitals $\psi_{\lambda}$, with $\lambda=1,2, \ldots, A$.

The unitary transformation

$$
\psi_{a}=\sum_{\lambda} C_{a \lambda} \phi_{\lambda},
$$

brings us into the new basis. The new basis has quantum numbers $a=1,2, \ldots, A$. aragraph!paragraph $>$ paragraph $>-0.5 \mathrm{em}$
a) Show that the new basis is orthogonal.
aragraph!paragraph $>$ paragraph $>-0.5 \mathrm{em}$
b) Show that the new Slater determinant constructed from the new singleparticle wave functions can be written as the determinant based on the previous basis and the determinant of the matrix $C$.
aragraph!paragraph $>$ paragraph $>-0.5 \mathrm{em}$
c) Show that the old and the new Slater determinants are equal up to a complex constant with absolute value unity.

Hint. Hint: $C$ is a unitary matrix.

## *

Exercise 2: Matrix elements for the Hartree-Fock method and the nuclear shell model

We will assume that we can build various Slater determinants using an orthogonal single-particle basis $\psi_{\lambda}$, with $\lambda=1,2, \ldots, A$.

The aim of this exercise is to set up specific matrix elements that will turn useful when we start our discussions of the nuclear shell model. In particular you will notice, depending on the character of the operator, that many matrix elements will actually be zero.

Consider three $A$-particle Slater determinants $\left.\left|\Phi_{0},\right| \Phi_{i}^{a}\right\rangle$ and $\left|\Phi_{i j}^{a b}\right\rangle$, where the notation means that Slater determinant $\left|\Phi_{i}^{a}\right\rangle$ differs from $\left|\Phi_{0}\right\rangle$ by one singleparticle state, that is a single-particle state $\psi_{i}$ is replaced by a single-particle state $\psi_{a}$. It will later be interpreted as a so-called one-particle-one-hole excitation. Similarly, the Slater determinant $\left|\Phi_{i j}^{a b}\right\rangle$ differs by two single-particle states from $\left|\Phi_{0}\right\rangle$ and is normally thought of as a two-particle-two-hole excitation.

Define a general onebody operator $\hat{F}=\sum_{i}^{A} \hat{f}\left(x_{i}\right)$ and a general twobody operator $\hat{G}=\sum_{i>j}^{A} \hat{g}\left(x_{i}, x_{j}\right)$ with $g$ being invariant under the interchange of the coordinates of particles $i$ and $j$. You can use here the results from the second exercise set, exercise 3 .
aragraph!paragraph $>$ paragraph $>-0.5 \mathrm{em}$
a) Calculate

$$
\left\langle\Phi_{0}\right| \hat{F}\left|\Phi_{0}\right\rangle
$$

and

$$
\begin{array}{r}
\left\langle\Phi_{0}\right| \hat{G}\left|\Phi_{0}\right\rangle . \\
\text { aragraph!paragraph }>\text { paragraph }>-0.5 \mathrm{em}
\end{array}
$$

b) Find thereafter

$$
\left\langle\Phi_{0}\right| \hat{F}\left|\Phi_{i}^{a}\right\rangle
$$

and

$$
\left\langle\Phi_{0}\right| \hat{G}\left|\Phi_{i}^{a}\right\rangle
$$

aragraph!paragraph $>$ paragraph $>-0.5 \mathrm{em}$
c) Finally, find

$$
\left\langle\Phi_{0}\right| \hat{F}\left|\Phi_{i j}^{a b}\right\rangle,
$$

and

$$
\left\langle\Phi_{0}\right| \hat{G}\left|\Phi_{i j}^{a b}\right\rangle .
$$

What happens with the two-body operator if we have a transition probability of the type

$$
\left\langle\Phi_{0}\right| \hat{G}\left|\Phi_{i j k}^{a b c}\right\rangle
$$

where the Slater determinant to the right of the operator differs by more than two single-particle states?
aragraph!paragraph $>$ paragraph $>-0.5 \mathrm{em}$
d) With an orthogonal basis of Slater determinants $\Phi_{\lambda}$, we can now construct an exact many-body state as a linear expansion of Slater determinants, that is, a given exact state

$$
\Psi_{i}=\sum_{\lambda=0}^{\infty} C_{i \lambda} \Phi_{\lambda}
$$

In all practical calculations the infinity is replaced by a given truncation in the sum.

If you are to compute the expectation value of (at most) a two-body Hamiltonian for the above exact state

$$
\left\langle\Psi_{i}\right| \hat{H}\left|\Psi_{i}\right\rangle
$$

based on the calculations above, which are the only elements which will contribute? (there is no need to perform any calculation here, use your results from exercises a), b), and c)).

These results simplify to a large extent shell-model calculations.

## *

## Project 3: Developing a Hartree-Fock program

Here we will demonstrate how do develop a Hartree-Fock program that will allow us to compute the binding energy of nuclei like ${ }^{16} \mathrm{O}$ or ${ }^{40} \mathrm{Ca}$. The singleparticle energies obtained by solving the Hartree-Fock equations can then be directly related to experimental separation energies. Since Hartree-Fock theory is the starting point for several many-body techniques (density functional theory, random-phase approximation, shell-model etc), the aim here is to develop a computer program to solve the Hartree-Fock equations in a given single-particle basis, here the harmonic oscillator.

The Hamiltonian for a system of $A$ nucleons confined in a harmonic potential reads

$$
\hat{H}=\sum_{i=1}^{A} \frac{\hat{p}_{i}^{2}}{2 m}+\sum_{i=1}^{A} \frac{1}{2} m \omega r_{i}^{2}+\sum_{i<j} \hat{V}_{i j}
$$

with $\hbar^{2} / 2 m=20.73 \mathrm{fm}^{2}, m c^{2}=938.926 \mathrm{MeV}$, and $\hat{V}_{i j}$ is the two-body interaction potential whose matrix elements are precalculated and to be read in by you, see weblink below.

The Hartree-Fock algorithm can be broken down as follows. We recall that our Hartree-Fock matrix is

$$
\hat{h}_{\alpha \beta}^{H F}=\langle\alpha| \hat{h}_{0}|\beta\rangle+\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{*} C_{j \delta}\langle\alpha \gamma| V|\beta \delta\rangle_{A S} .
$$

Normally we assume that the single-particle basis $|\beta\rangle$ forms an eigenbasis for the operator $\hat{h}_{0}$ (this is our case), meaning that the Hartree-Fock matrix becomes

$$
\hat{h}_{\alpha \beta}^{H F}=\epsilon_{\alpha} \delta_{\alpha, \beta}+\sum_{j=1}^{A} \sum_{\gamma \delta} C_{j \gamma}^{*} C_{j \delta}\langle\alpha \gamma| V|\beta \delta\rangle_{A S} .
$$

The Hartree-Fock eigenvalue problem

$$
\sum_{\beta} \hat{h}_{\alpha \beta}^{H F} C_{i \beta}=\epsilon_{i}^{\mathrm{HF}} C_{i \alpha},
$$

can be written out in a more compact form as

$$
\hat{h}^{H F} \hat{C}=\epsilon^{\mathrm{HF}} \hat{C}
$$

The equations are often rewritten in terms of a so-called density matrix, which is defined as

$$
\begin{equation*}
\rho_{\gamma \delta}=\sum_{i=1}^{A}\langle\gamma \mid i\rangle\langle i \mid \delta\rangle=\sum_{i=1}^{A} C_{i \gamma} C_{i \delta}^{*} . \tag{18}
\end{equation*}
$$

It means that we can rewrite the Hartree-Fock Hamiltonian as

$$
\hat{h}_{\alpha \beta}^{H F}=\epsilon_{\alpha} \delta_{\alpha, \beta}+\sum_{\gamma \delta} \rho_{\gamma \delta}\langle\alpha \gamma| V|\beta \delta\rangle_{A S} .
$$

It is convenient to use the density matrix since we can precalculate in every iteration the product of two eigenvector components $C$.

Note that $\langle\alpha| \hat{h}_{0}|\beta\rangle$ denotes the matrix elements of the one-body part of the starting hamiltonian. For self-bound nuclei $\langle\alpha| \hat{h}_{0}|\beta\rangle$ is the kinetic energy while $\langle\alpha| \hat{h}_{0}|\beta\rangle$ represents the harmonic oscillator hamiltonian since the system is confined in a harmonic trap. If we are working in a harmonic oscillator basis with the same $\omega$ as the trapping potential, then $\langle\alpha| \hat{h}_{0}|\beta\rangle$ is diagonal.

An example of a function written in python which performs the Hartree-Fock calculation is shown here. In setting up your code you will need to write a function which sets up the single-particle basis, the matrix elements $t_{\alpha \gamma}$ of the one-body operator (called $h 0$ in the function below) and the antisymmetrized TBMEs (called nninteraction in the code link below) and the density matrix elements $\rho_{\beta \delta}$ (called densityMatrix below). The python program shows how one can, in a brute force way read in matrix elements in $m$-scheme and compute the Hartree-Fock single-particle energies for four major shells. The interaction which has been used is the so-called N3LO interaction of Machleidt and Entem using the Similarity Renormalization Group approach method to renormalize the interaction, using an oscillator energy $\hbar \omega=10 \mathrm{MeV}$.

The codes used to obtain the matrix elements are available at the CENS site. The python program shows how one can, in a brute force way read in matrix elements in $m$-scheme and compute the Hartree-Fock single-particle energies
for four major shells. The interaction which has been used is the so-called N3LO interaction of Machleidt and Entem using the Similarity Renormalization Group approach method to renormalize the interaction, using an oscillator energy $\hbar \omega=10 \mathrm{MeV}$.

The nucleon-nucleon two-body matrix elements are in $m$-scheme and are fully anti-symmetrized. The Hartree-Fock programs uses the density matrix discussed above in order to compute the Hartree-Fock matrix. Here we display the Hartree-Fock part only, assuming that single-particle data and two-body matrix elements have already been read in.

```
""" Star HF-iterations, preparing variables and density matrix """
""" Coefficients for setting up density matrix, assuming only one along the diagonals """
C = np.eye(spOrbitals) # HF coefficients
DensityMatrix = np.zeros([spOrbitals,spOrbitals])
for gamma in range(spOrbitals):
    for delta in range(spOrbitals):
            sum = 0.0
            for i in range(Nparticles):
            sum += C[gamma][i]*C[delta] [i]
            DensityMatrix[gamma][delta] = Decimal(sum)
maxHFiter = 100
epsilon = 1.0e-10
difference = 1.0
hf_count = 0
oldenergies = np.zeros(spOrbitals)
newenergies = np.zeros(spOrbitals)
while hf_count < maxHFiter and difference > epsilon:
    print "############### Iteration %i ###############" % hf_count
            HFmatrix = np.zeros([spOrbitals,spOrbitals])
    for alpha in range(spOrbitals):
                for beta in range(spOrbitals):
                    """ If tests for three-dimensional systems, including isospin conser
                    if l[alpha] != l[beta] and j[alpha] != j[beta] and mj[alpha] != mj[be
                        """ Setting up the Fock matrix using the density matrix and antisymm
                        sumFockTerm = 0.0
                    for gamma in range(spOrbitals):
                    for delta in range(spOrbitals):
                    if (mj[alpha]+mj[gamma]) != (mj[beta]+mj[delta]) and (tz[alph
                    sumFockTerm += DensityMatrix[gamma] [delta]*nninteraction[alph
                    HFmatrix[alpha][beta] = Decimal(sumFockTerm)
                    """ Adding the one-body term, here plain harmonic oscillator """
                    if beta == alpha: HFmatrix[alpha][alpha] += singleparticleH[alpha]
    spenergies, C = np.linalg.eigh(HFmatrix)
    """ Setting up new density matrix in m-scheme """
    DensityMatrix = np.zeros([spOrbitals,spOrbitals])
    for gamma in range(spOrbitals):
            for delta in range(spOrbitals):
                    sum = 0.0
                    for i in range(Nparticles):
                    sum += C[gamma][i]*C[delta][i]
            DensityMatrix[gamma][delta] = Decimal(sum)
    newenergies = spenergies
    """ Brute force computation of difference between previous and new sp HF energies
    sum =0.0
    for i in range(spOrbitals):
            sum += (abs(newenergies[i]-oldenergies[i]))/spOrbitals
    difference = sum
```

```
oldenergies = newenergies
print "Single-particle energies, ordering may have changed "
for i in range(spOrbitals):
    print('{0:4d} {1:.4f}'.format(i, Decimal(oldenergies[i])))
hf_count += 1
```

Running the program, one finds that the lowest-lying states for a nucleus like ${ }^{16} \mathrm{O}$, we see that the nucleon-nucleon force brings a natural spin-orbit splitting for the $0 p$ states (or other states except the $s$-states). Since we are using the $m$-scheme for our calculations, we observe that there are several states with the same eigenvalues. The number of eigenvalues corresponds to the degeneracy $2 j+1$ and is well respected in our calculations, as see from the table here.

The values of the lowest-lying states are ( $\pi$ for protons and $\nu$ for neutrons)

| Quantum numbers | Energy $[\mathrm{MeV}]$ |
| :---: | :---: |
| $0 s_{1 / 2}^{\pi}$ | -40.4602 |
| $0 s_{1 / 2}^{\pi}$ | -40.4602 |
| $0 s_{1 / 2}^{\nu}$ | -40.6426 |
| $0 s_{1 / 2}^{\nu}$ | -40.6426 |
| $0 p_{1 / 2}^{\pi}$ | -6.7133 |
| $0 p_{1 / 2}^{\pi}$ | -6.7133 |
| $0 p_{1 / 2}^{\nu}$ | -6.8403 |
| $0 p_{1 / 2}^{\nu}$ | -6.8403 |
| $0 p_{3 / 2}^{\pi}$ | -11.5886 |
| $0 p_{3 / 2}^{\pi}$ | -11.5886 |
| $0 p_{3 / 2}^{\pi}$ | -11.5886 |
| $0 p_{3 / 2}^{\pi}$ | -11.5886 |
| $0 p_{3 / 2}^{\nu}$ | -11.7201 |
| $0 p_{3 / 2}^{\nu}$ | -11.7201 |
| $0 p_{3 / 2}^{\nu}$ | -11.7201 |
| $0 p_{3 / 2}^{\nu}$ | -11.7201 |
| $0 d_{5 / 2}^{\pi}$ | 18.7589 |
| $0 d_{5 / 2}^{\nu}$ | 18.8082 |

We can use these results to attempt our first link with experimental data, namely to compute the shell gap or the separation energies. The shell gap for neutrons is given by

$$
\Delta S_{n}=2 B E(N, Z)-B E(N-1, Z)-B E(N+1, Z)
$$

For ${ }^{16} \mathrm{O}$ we have an experimental value for the shell gap of 11.51 MeV for neutrons, while our Hartree-Fock calculations result in 25.65 MeV . This means that correlations beyond a simple Hartree-Fock calculation with a two-body force play an important role in nuclear physics. The splitting between the $0 p_{3 / 2}^{\nu}$ and the $0 p_{1 / 2}^{\nu}$ state is 4.88 MeV , while the experimental value for the gap between the ground state $1 / 2^{-}$and the first excited $3 / 2^{-}$states is 6.08 MeV . The two-nucleon
spin-orbit force plays a central role here. In our discussion of nuclear forces we will see how the spin-orbit force comes into play here.

